

Testing SUSY Unification

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Two Scale Picture of Nature

*Electroweak
Scale
 $\langle H \rangle \simeq 10^2 \text{ GeV}$*



*GUT/String/Planck
Scale
 10^{16-19} GeV*

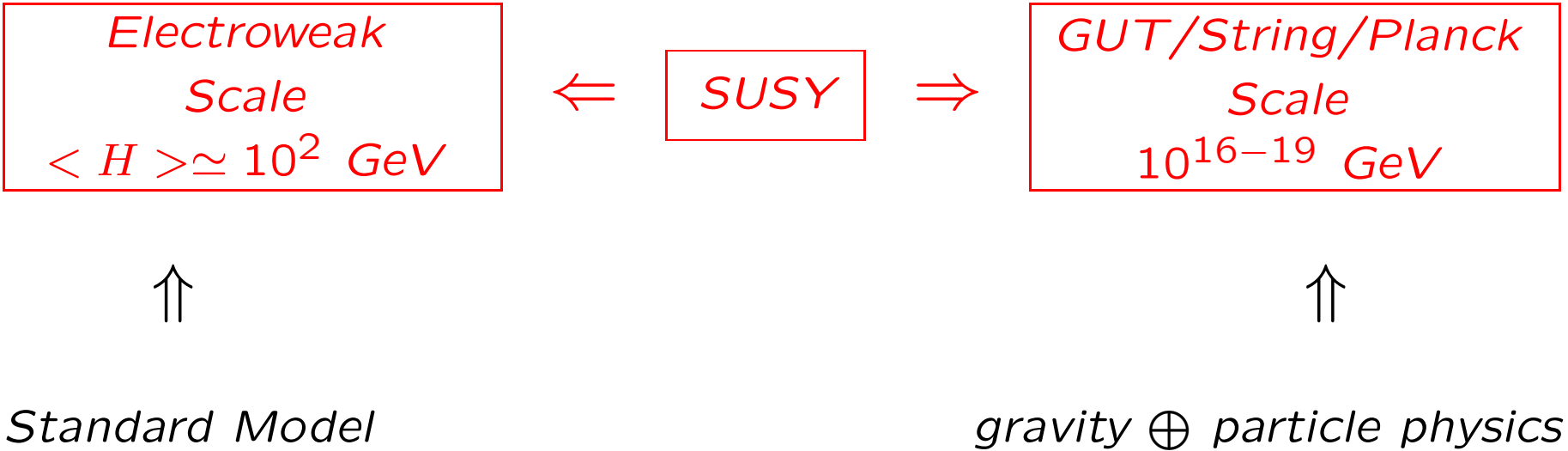


Standard Model



gravity \oplus particle physics

Two Scale Picture of Nature



SUSY: allows also for gauge coupling unification, radiative electroweak symmetry breaking, ...

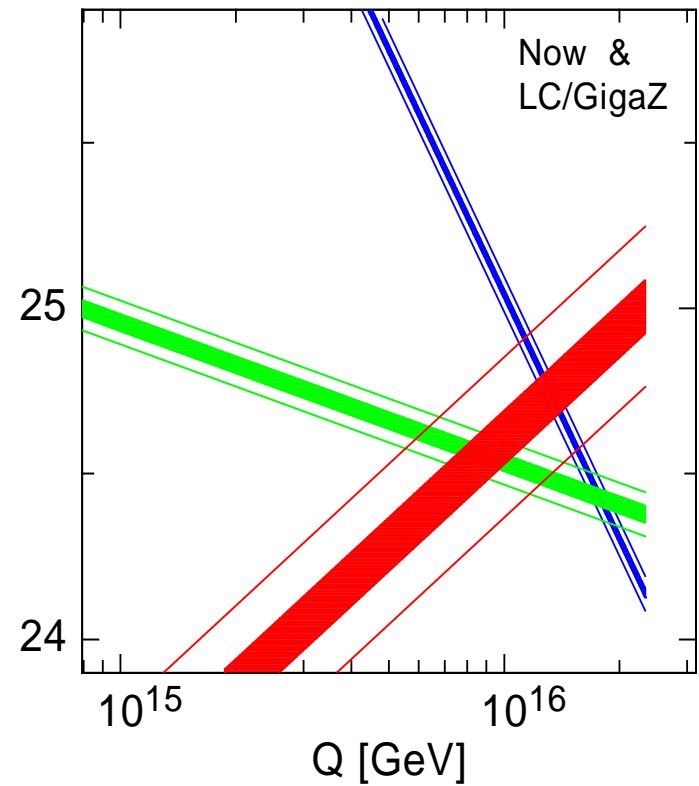
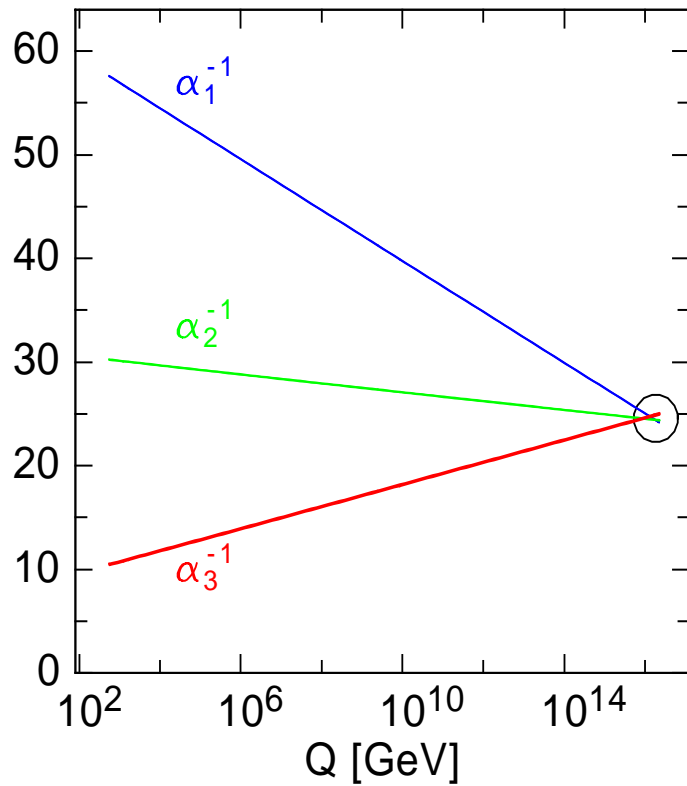
Exploring high scale structures (GUT, PL ...)

- *Proton decay*
- *Cosmology at early time of the universe*
- *Neutrino physics (see-saw), fermion mass textures*
- *Extrapolation of high precision parameters:*
 - gauge and Yukawa couplings*
 - SUSY parameters*

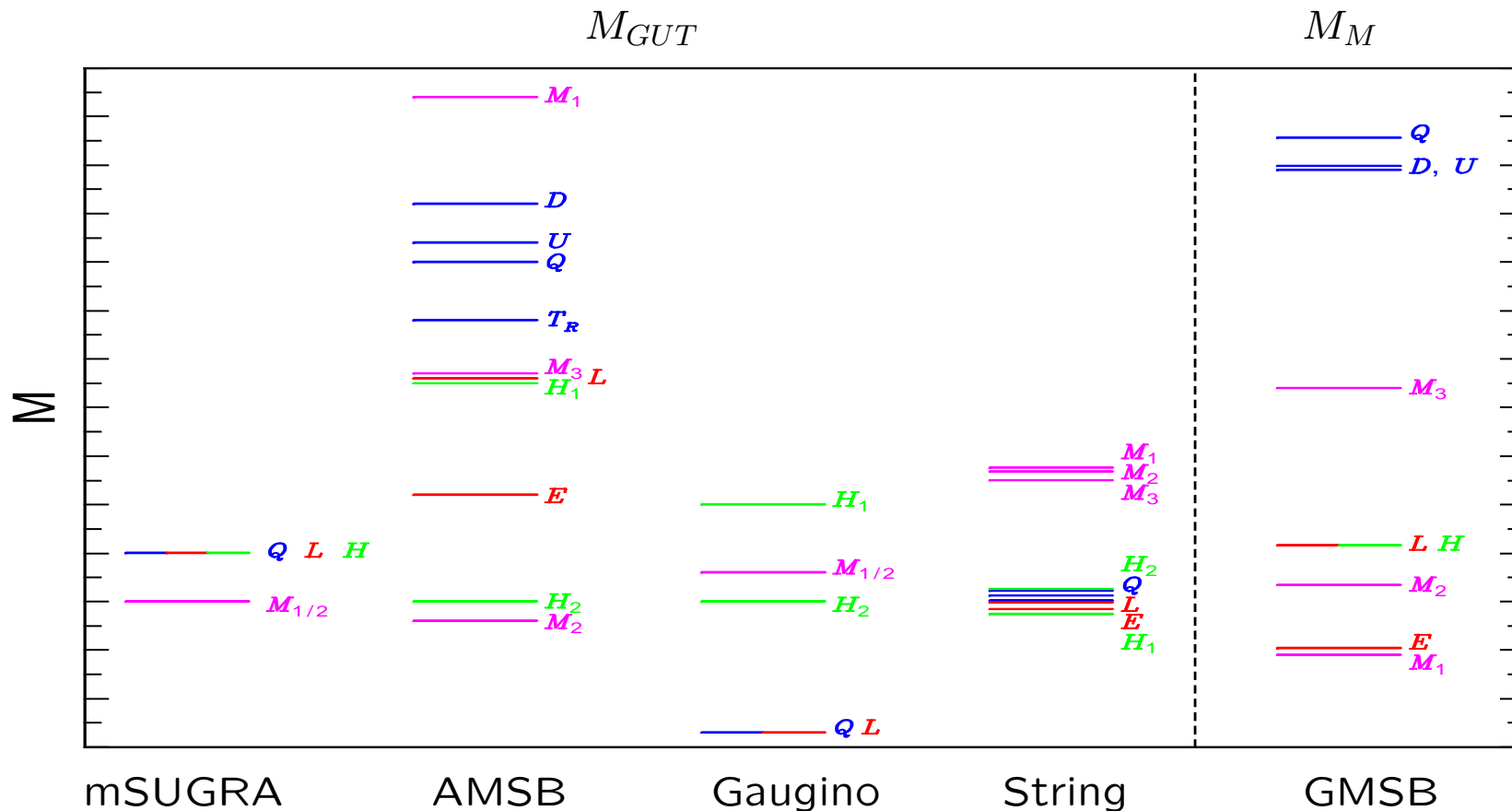
Experimental information

- *LEP/Tevatron:*
Higgs heavier than 100 GeV
charginos/sleptons heavier than 100 GeV
squarks (except \tilde{t}, \tilde{b}), gluinos heavier than 200 GeV
- *rare decays:*
bounds on flavour violation beyond CKM
- *Cold dark matter: $\Omega h^2 \lesssim 0.11$*
- *high precision measurements of gauge couplings*
 \Rightarrow unification if SUSY is present

Evolution of gauge couplings



Regularities at High Scales



Low Energy Parameters

Measurements:

masses
cross sections
polarization



SUSY parameters:

gaugino parameters M_i
scalar masses: $M_{H_i}^2, M_E^2, M_L^2, \dots$
Higgs/Higgsino parameters: $\mu, \tan \beta$
trilinear couplings: A_t, A_b, A_τ

Mass measurements, LHC

SPS1a (bulk region)

$m_0 = 100$ GeV,

$m_{1/2} = 250$ GeV,

$A_0 = -100$ GeV,

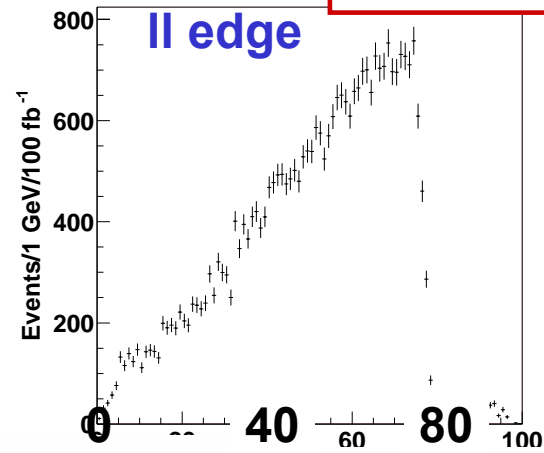
$\tan(\beta) = 10, \mu > 0$

Left squark cascade decay

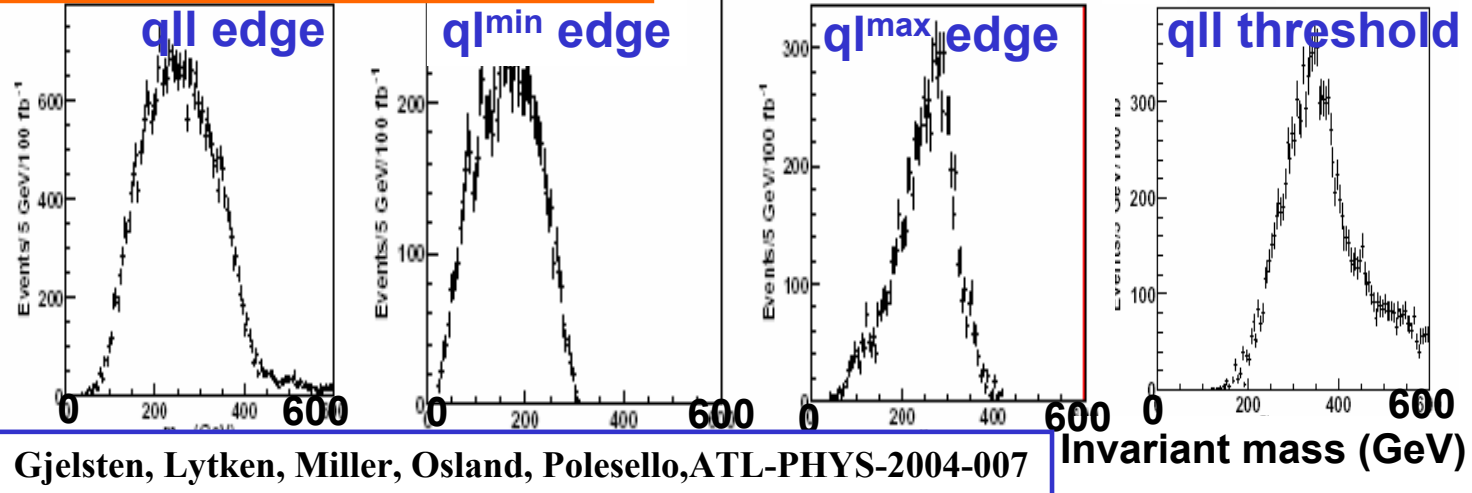
$$\tilde{q}_L \rightarrow \tilde{\chi}_2^0 q \rightarrow \tilde{l}_R l q \rightarrow llq \tilde{\chi}_1^0$$

fast sim.

$L = 100 \text{ fb}^{-1}$



2 SFOS lep., $p_T > 20, 10$ GeV
 ≥ 4 jets, $p_T > 150, 100, 50, 50$ GeV
 $M_{\text{eff}} > 600$ GeV
 $E_{T\text{miss}} > \max(100, 0.2 M_{\text{eff}})$



Gjelsten, Lykken, Miller, Osland, Polesello, ATL-PHYS-2004-007

Mass measurements, LHC

$L=100 \text{ fb}^{-1}$

Fit results

Edge	Nominal Value	Fit Value	Syst. Error Energy Scale	Statistical Error
$m(ll)^{\text{edge}}$	77.077	77.024	0.08	0.05
$m(qll)^{\text{edge}}$	431.1	431.3	4.3	2.4
$m(ql)_{\text{min}}^{\text{edge}}$	302.1	300.8	3.0	1.5
$m(ql)_{\text{max}}^{\text{edge}}$	380.3	379.4	3.8	1.8
$m(qll)^{\text{thres}}$	203.0	204.6	2.0	2.8

Mass reconstruction

5 endpoints measurements, 4 unknown masses

$$\chi^2 = \sum \chi_j^2 = \sum \left[\frac{E_j^{\text{theory}}(\vec{m}) - E_j^{\text{exp}}}{\sigma_j^{\text{exp}}} \right]^2$$

$$E_j^i = E_j^{\text{nom}} + a_j^i \sigma_j^{\text{fit}} + b_j^i \sigma_j^{\text{scale}}$$

$$m(\chi_1^0) = 96 \text{ GeV}$$

$$m(l_R) = 143 \text{ GeV}$$

$$m(\chi_2^0) = 177 \text{ GeV}$$

$$m(q_L) = 540 \text{ GeV}$$

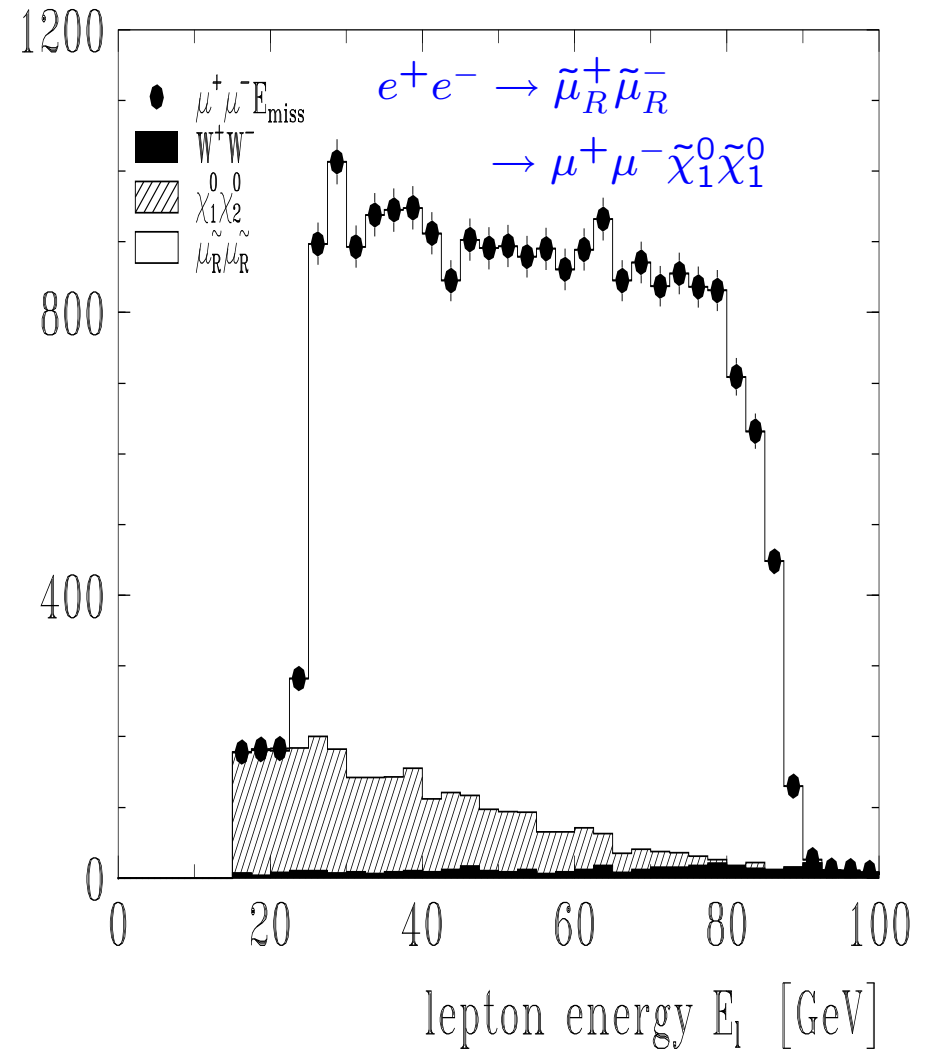
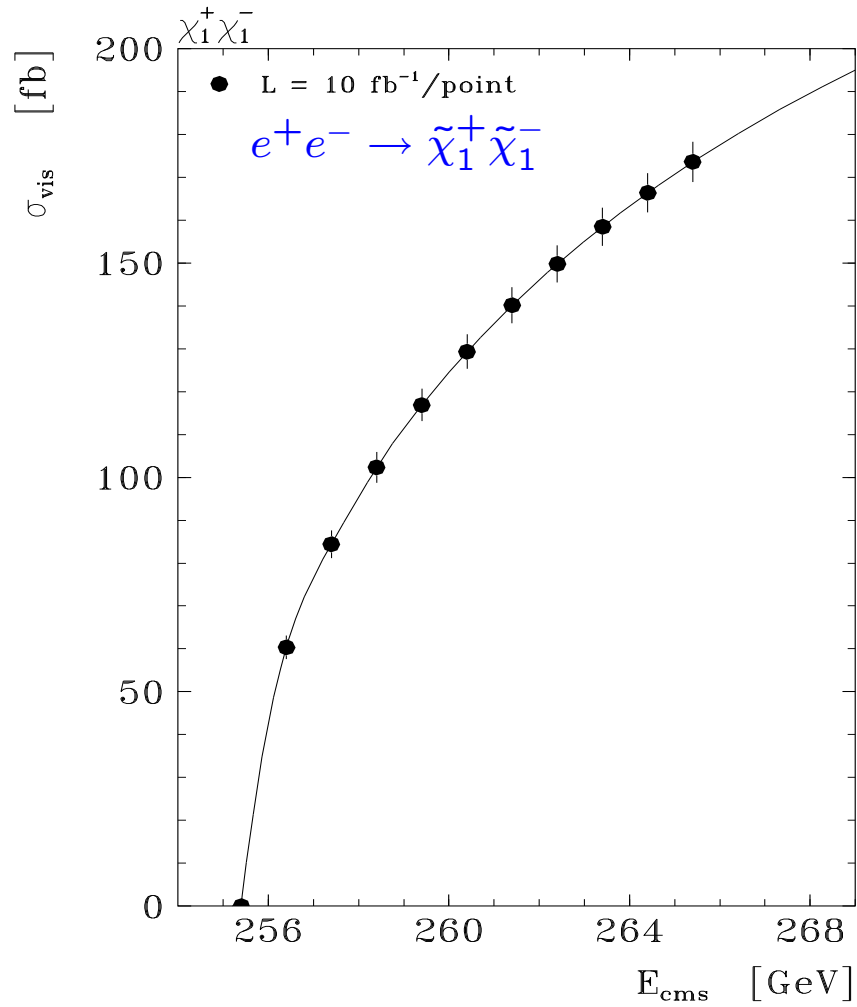
$$\Delta m(\chi_1^0) = 4.8 \text{ GeV}, \quad \Delta m(\chi_2^0) = 4.7 \text{ GeV},$$

$$\Delta m(l_R) = 4.8 \text{ GeV}, \quad \Delta m(q_L) = 8.7 \text{ GeV}$$

Gjelsten, Lytken, Miller, Osland, Polesello, ATL-PHYS-2004-007

Mass measurements, ILC

U. Martyn



Expected Accuracies

LHC: masses of squarks, gluinos, winos, bino within a few per-cent

LC: sleptons, winos, bino within per-mile

typical values for mSUGRA scenario

$\tilde{\chi}_1^+$	183.05 ± 0.15	0.08 %	\tilde{e}_R	224.82 ± 0.15	0.06 %
$\tilde{\chi}_2^+$	385.28 ± 0.28		\tilde{e}_L	269.09 ± 0.28	
$\tilde{\chi}_1^0$	97.86 ± 0.20	0.2 %	\tilde{u}_R	572.0 ± 10.0	1.8 %
$\tilde{\chi}_2^0$	184.65 ± 0.30		\tilde{u}_L	589.0 ± 10.0	

LHC + LC: combining data of both machines can improve accuracies on some masses considerably, e.g. $\Delta m_{\tilde{\chi}_2^0}$ up to an order of magnitude. (B.K. Gjelsten, D. Miller, P. Osland and G. Polesello)

RGE structures

explicit solutions of 1-loop RGEs:

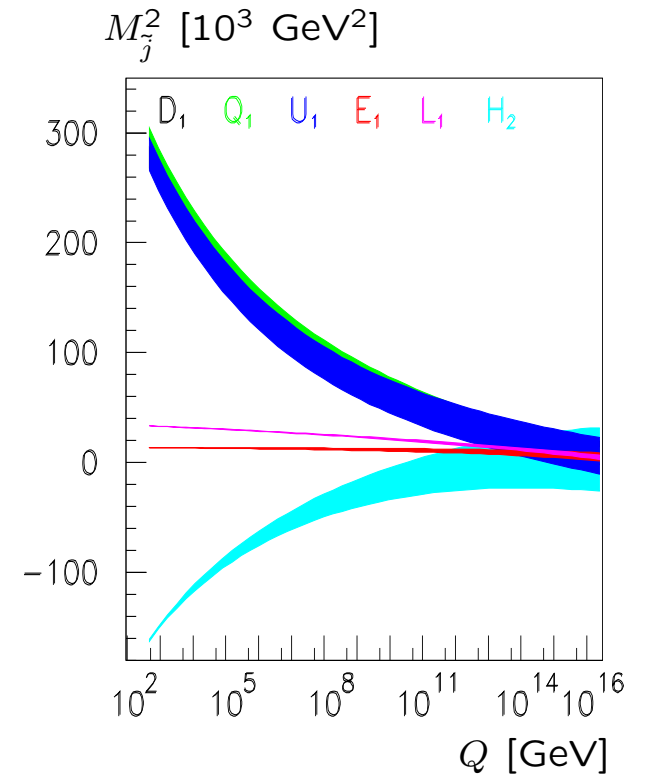
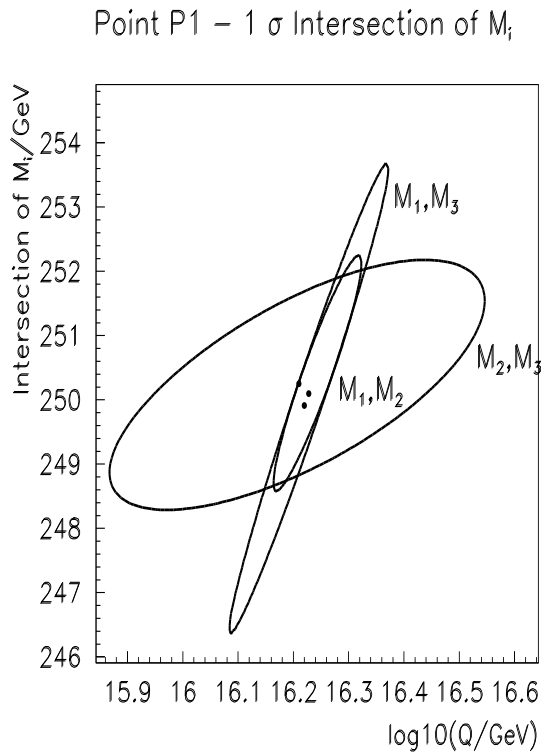
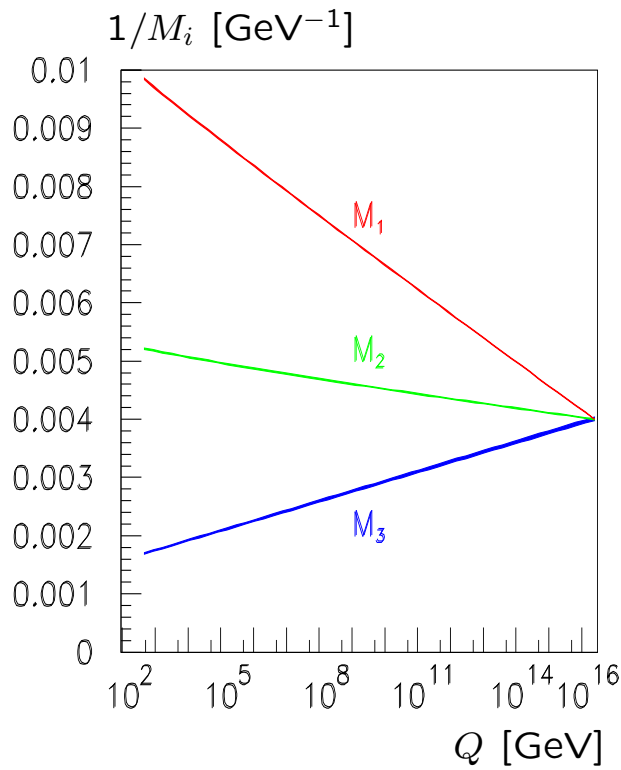
$$\begin{aligned} M_1 &= 0.41 M_{1/2} && \Rightarrow M_{1/2} \text{ easy} \\ M_L^2 &= M_0^2 + 0.47 M_{1/2}^2 && \Rightarrow M_0 \text{ easy} \\ M_Q^2 &= M_0^2 + 5.1 M_{1/2}^2 && \Rightarrow M_0 \text{ difficult} \\ M_{H_2}^2 &= -0.03 M_0^2 - 1.34 M_{1/2}^2 + \dots && \Rightarrow M_0 \text{ very difficult} \end{aligned}$$

Top-Down (taking *mSUGRA* as example)

$$\begin{aligned} M_{1/2} &= 250 \pm 0.08 \text{ GeV} \\ M_0 &= 200 \pm 0.09 \text{ GeV} \\ A_0 &= -100 \pm 1.8 \text{ GeV} \end{aligned}$$

mSUGRA

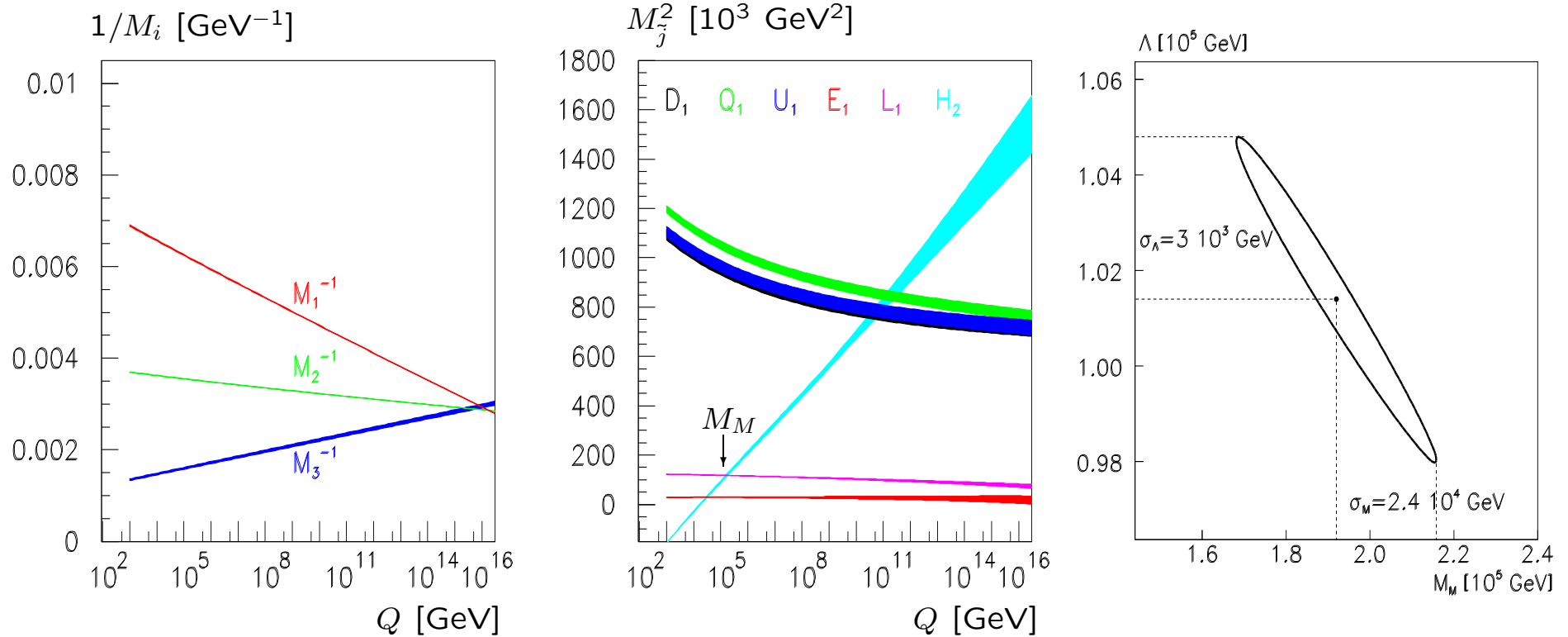
$\tan \beta = 10, M_0 = 70 \text{ GeV}, M_{1/2} = 250 \text{ GeV}, A_0 = -300, \text{sign}(\mu) = 1$



1 σ error bands

GMSB

$M_M = 200 \text{ TeV}$, $\Lambda = 100 \text{ TeV}$, $N_5 = 1$, $\tan \beta = 15$, $A_0 = 0$, $\text{sign}(\mu) = 1$



1 σ error bands

Summary

- *Reconstruction of the underlying high scale theory is feasible*
- *– LHC allows for measurement in the per-cent range*
 - most likely information on spectrum will not be complete*
 - ⇒ top-down fits to exclude models*
- *High precision measurements at future e^+e^- colliders are necessary for bottom-up approach*